

Radiation

Simple Radiating Systems

Consider a localized source oscillating at frequency ω :

$$\rho(\vec{x}, t) = \rho(\vec{x}) e^{-i\omega t}, \quad \vec{J}(\vec{x}, t) = \vec{J}(\vec{x}) e^{-i\omega t}$$

In the Lorenz gauge (which is good for localized sources), we have:

$$\left(\nabla^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2}\right) \vec{A} = -\mu_0 \vec{J}$$

Then (since there are no sources at ∞):

$$\vec{A}(\vec{x}, t) = \frac{\mu_0}{4\pi} \int \frac{\vec{J}(\vec{x}', t') \delta(t - t' - \frac{|\vec{x} - \vec{x}'|}{c})}{|\vec{x} - \vec{x}'|} d^3x' dt'$$

← retarded Green's function

Integration over t' results in:

$$\vec{A}(\vec{x}, t) = \frac{\mu_0}{4\pi} \int \frac{\vec{J}(\vec{x}', t - \frac{|\vec{x} - \vec{x}'|}{c})}{|\vec{x} - \vec{x}'|} d^3x'$$

After using $\vec{J}(\vec{x}, t) = \vec{J}(\vec{x}) e^{-i\omega t}$, we find:

$$\vec{A}(\vec{x}, t) = \vec{A}(\vec{x}) e^{-i\omega t}$$

Where:

$$\vec{A}(\vec{x}) = \frac{\mu_0}{4\pi} \int \vec{J}(\vec{x}') \frac{e^{i\frac{\omega}{c}|\vec{x}-\vec{x}'|}}{|\vec{x}-\vec{x}'|} d^3x'$$

We consider the long wavelength limit where the source dimension

"d" is small compared to the wavelength $\lambda \equiv \frac{2\pi c}{\omega}$. The three

possible cases then are:

(a) Near-field region ($d \ll \lambda, r \ll \lambda$). In this case $\frac{\omega}{c} |\vec{x}-\vec{x}'| \ll 2\pi$,

and hence:

$$\vec{A}(\vec{x}) \approx \frac{\mu_0}{4\pi} \int \frac{\vec{J}(\vec{x}')}{|\vec{x}-\vec{x}'|} d^3x'$$

This is called the "static limit", in which (apart from the $e^{-i\omega t}$

factor) \vec{A} is the same as that for a steady current. Therefore,

in this case, magnetostatics holds.

(b) Far-field region ($d \ll \lambda, r \gg d, \lambda$). In this case, we have:

$$\begin{aligned} \vec{A}(\vec{x}) &\approx \frac{\mu_0}{4\pi} \int \frac{\vec{J}(\vec{x}')}{r} e^{ik\sqrt{r^2 + r'^2 - 2\vec{x} \cdot \vec{x}'}} d^3x' \approx \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \int \vec{J}(\vec{x}') \\ &e^{-ik\frac{\vec{x} \cdot \vec{x}'}{r}} d^3x' \Rightarrow A(\vec{x}) \approx \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \int \vec{J}(\vec{x}') e^{-ik\hat{n} \cdot \vec{x}'} d^3x' \\ &\quad \downarrow \\ &\quad h \equiv \frac{\vec{x}}{r} \end{aligned}$$

The integral is essentially the Fourier transform of the current density.

(c) Arbitrary observation distance. In this case, we use:

$$\frac{e^{ik|\vec{x}-\vec{x}'|}}{|\vec{x}-\vec{x}'|} = ik \sum_{l,m} j_l(kr_<) h_l^{(1)}(kr_>) Y_{lm}(\theta, \phi) Y_{lm}^*(\theta', \phi')$$

Here $j_l(x) = (-x)^l \left(\frac{1}{x} \frac{d}{dx}\right)^l \left(\frac{\sin x}{x}\right)$ are the spherical Bessel functions.

Also:

$$h_l^{(1)}(x) = j_l(x) + in_l(x)$$

Where:

$$n_l(x) = (-x)^l \left(\frac{1}{x} \frac{d}{dx}\right)^l \left(\frac{-\cos x}{x}\right)$$

For \vec{x} sufficiently outside the source, $r_< = r'$ and $r_> = r$. Hence:

$$\vec{A}(\vec{x}) = \mu_0 ik \sum_{l,m} h_l^{(1)}(kr) Y_{lm}(\theta, \phi) \left[\int \vec{J}(\vec{x}') j_l(kr') Y_{lm}^*(\theta', \phi') d^3x' \right]$$

In the long wavelength limit, $kr' \ll 2\pi$ and $j_l(kr') \approx \frac{(kr')^l}{(2l+1)!!}$,

where $(2l+1)!! = (2l+1)(2l-1)\dots 3 \cdot 1$.

then,

keeping only the first term $l=0$, we have $j_0(kr') \approx 1$ and

$$h^{(1)}(kr) = \frac{e^{ikr}}{ikr}. \text{ This results in:}$$

$$\vec{A}(\vec{x}, t) \approx \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \int \vec{J}(\vec{x}', t) d^3x' = \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \int \vec{\nabla}' \cdot (\vec{x}' \vec{J}) d^3x'$$

(localized source)

$$-\frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \int \vec{x}' \vec{\nabla}' \cdot \vec{J} d^3x'$$

Unlike magnetostatic, where $\vec{\nabla}' \cdot \vec{J}(\vec{x}') = 0$, we now have (from conservation of electric charge):

$$\vec{\nabla}' \cdot \vec{J}(\vec{x}', t) + \frac{\partial \rho(\vec{x}', t)}{\partial t} = 0 \Rightarrow \vec{\nabla}' \cdot \vec{J}(\vec{x}') = i\omega \rho(\vec{x}')$$

Thus:

$$\vec{A}(\vec{x}, t) \approx -i\omega \int \vec{x}' \rho(\vec{x}') d^3x' \Rightarrow \vec{A}(\vec{x}, t) = \frac{-i\mu_0\omega}{4\pi} \frac{\vec{P}(\vec{x})}{r} e^{ikr} e^{-i\omega t}$$

Here $\vec{P} e^{-i\omega t} = \int \vec{x}' \rho(\vec{x}', t) d^3x'$ is the oscillating electric dipole moment of the source.

Therefore:

$$\vec{B}(\vec{x}, t) = \vec{\nabla} \times \vec{A}(\vec{x}, t) = \frac{-i\mu_0\omega}{4\pi} \vec{\nabla} \left(\frac{e^{ikr}}{r} \right) \times \vec{P} e^{-i\omega t} = \frac{\mu_0\omega k}{4\pi} \frac{e^{ikr}}{r} \left(1 - \frac{1}{ikr}\right) \hat{r} \times \vec{P} e^{-i\omega t}$$

Also, from Maxwell equations, we have:

$$-i\omega \vec{E}(\vec{x}, t) = \frac{1}{\mu_0 \epsilon_0} \vec{\nabla} \times \vec{B}(\vec{x}, t) \Rightarrow \vec{E}(\vec{x}, t) = \frac{i}{\mu_0 \epsilon_0 \omega} \vec{\nabla} \times \vec{B}(\vec{x}, t)$$

Hence:

$$\vec{E}(\vec{x}, t) = \frac{ik}{4\pi} \vec{\nabla} \times \left[\frac{e^{ikr}}{r} \left(1 - \frac{1}{ikr}\right) \hat{n} \times \vec{p} \right] e^{-i\omega t}$$

Note that $\hat{n} \equiv \frac{\vec{x}}{r}$. Then:

$$\vec{\nabla} \times \left[\frac{e^{ikr}}{r} \left(1 - \frac{1}{ikr}\right) \hat{n} \times \vec{p} \right] = \vec{\nabla} \times \left[\frac{e^{ikr}}{r} \left(1 - \frac{1}{ikr}\right) \frac{\vec{x}}{r} \times \vec{p} \right] =$$

$$\vec{\nabla} \left[\frac{e^{ikr}}{r^2} \left(1 - \frac{1}{ikr}\right) \right] \times (\vec{x} \times \vec{p}) + \frac{e^{ikr}}{r^2} \left(1 - \frac{1}{ikr}\right) \vec{\nabla} \times (\vec{x} \times \vec{p}) =$$

$$\left[\frac{e^{ikr}}{r^2} ik \left(1 - \frac{1}{ikr}\right) + e^{ikr} \left(-\frac{2}{r^3} + \frac{3}{ikr^4}\right) \right] \hat{n} \times (\hat{n} \times \vec{p}) r + \frac{e^{ikr}}{r^2}$$

$$\left(1 - \frac{1}{ikr}\right) \left[\underbrace{(\vec{p} \cdot \vec{\nabla}) \vec{x}}_{\vec{p}} - \underbrace{(\vec{\nabla} \cdot \vec{x}) \vec{p}}_3 \right] = ik \frac{e^{ikr}}{r} \hat{n} \times (\hat{n} \times \vec{p}) + e^{ikr}$$

$$\left(-\frac{3}{r^2} + \frac{3}{ikr^3}\right) \hat{n} \times (\hat{n} \times \vec{p}) - 2\vec{p} e^{ikr} \left(\frac{1}{r^2} - \frac{1}{ikr^3}\right) = ik \frac{e^{ikr}}{r} \hat{n} \times (\hat{n} \times \vec{p})$$

$$- e^{ikr} \left(\frac{1}{r^2} - \frac{1}{ikr^3}\right) \left[3(\hat{n} \cdot \vec{p}) \hat{n} - \vec{p} \right]$$

This leads to:

$$\vec{E}(\vec{x}, t) = \left[\frac{-k^2}{4\pi \epsilon_0} \frac{e^{ikr}}{r} \hat{n} \times (\hat{n} \times \vec{p}) + \frac{1}{4\pi \epsilon_0} (3\hat{n}(\hat{n} \cdot \vec{p}) - \vec{p}) \left(\frac{1}{r^3} - \frac{ik}{r^2}\right) \right] e^{-i\omega t}$$

In the near zone (i.e., $kr \ll 1$), we have:

$$\vec{E}(\vec{x}, t) \approx \frac{1}{4\pi\epsilon_0} \frac{3\hat{n}(\hat{n} \cdot \vec{p}) - \vec{p}}{r^3} e^{-i\omega t}$$

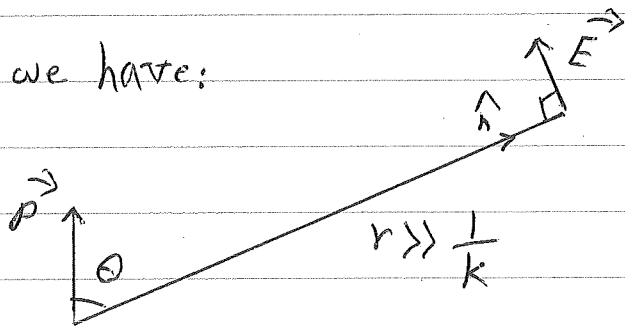
This is a quasi-static \vec{E} field. In the far zone (i.e., $kr \gg 1$):

$$\vec{E}(\vec{x}, t) \approx \frac{k^2}{4\pi\epsilon_0} \frac{e^{ikr}}{r} (\hat{n} \times \vec{p}) \times \hat{n}$$

It is important to note that $\vec{E} \perp \hat{n}$ in the far zone, while

$\vec{B} \perp \hat{n}$ in general (see the expression for \vec{B} on page (156)).

For a linearly polarized source, we have:



The time-averaged power per unit solid angle is given by:

$$\frac{dP}{d\Omega} = \text{Re} \left[\frac{1}{2} (\vec{E} \times \vec{H}^*) \cdot \hat{n} \right] r^2 = \frac{1}{2} \text{Re} \left[\vec{E} \cdot (\vec{H}^* \times \hat{n}) \right] r^2 \Rightarrow \frac{dP}{d\Omega} =$$

$$\frac{1}{2} \frac{ck^4}{(4\pi)^2 \epsilon_0} |(\hat{n} \times \vec{p}) \times \hat{n}|^2$$

For \vec{p} real, we have:

$$|(\hat{n} \times \vec{p}) \times \hat{n}| = |\hat{n} \times \vec{p}| = |\vec{p}| \sin \theta$$

Thus:

$$\frac{dP}{d\Omega} = \frac{ck^2}{32\pi^2\epsilon_0} |\vec{P}|^2 \sin^2\theta$$

The total radiated power is then given by:

$$P = \int \frac{ck^4}{32\pi^2\epsilon_0} |\vec{P}|^2 \sin^2\theta d\Omega = \frac{ck^4}{12\pi\epsilon_0} |\vec{P}|^2$$

Some comments are in order at this point. First, the expression for the radiated power is valid for a general polarization when the components of \vec{P} do not have the same phase. Second, the radiated power scales with the frequency as ω^4 . This is the Rayleigh scattering.

Example: A charge q oscillating at frequency ω along the z axis.

In this case, we have:

$$z(t) = z_0 \sin\omega t \quad (kz_0 \ll 1 \text{ in the long wavelength limit})$$

The physical electric dipole moment is:

$$\vec{P}_{\text{phys}} = qz(t)\hat{z} = qz_0 \sin\omega t \hat{z} = \text{Re}\left(\frac{iqz_0}{2} \hat{z} e^{-i\omega t}\right)$$

Hence:

$$\vec{P} = \frac{iqz_0}{2} \hat{z} \quad (\text{Linear polarization})$$

And:

$$\frac{dP}{d\Omega} = \frac{ck^4}{32\pi^2\epsilon_0} |\hat{n} \times \vec{P}|^2 = \frac{ck^4}{128\pi^2\epsilon_0} q^2 z_0^2$$

We note that in order to calculate \vec{P} we can calculate \vec{P}_{phys} first and then express it as the real part of $\vec{P}e^{-i\omega t}$.